Physics Beyond the Standard Model

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Beyond the Standard Model

Lecture I
•Why do we need to go Beyond the SM ?
•The Hierarchy Problem: what do we need to solve it ?

Lecture 2 • Supersymmetry and the Hierarchy Problem

Lecture 3 New Dynamics at the TeV scale: the Higgs as a (pseudo) Nambu-Goldstone Boson

Beyond the Standard Model I

•Status of the SM

•Why do we want to go beyond the SM ?

•A gauge theory: $SU(3)_c \times SU(2)_L \times U(1)_Y$

$$\begin{pmatrix} u \\ d \end{pmatrix}_{L}, \quad u_{R}, \quad d_{R} \quad \begin{pmatrix} \nu_{e} \\ e^{-} \end{pmatrix}_{L}, \quad e_{R}^{-}$$

$$\begin{pmatrix} c \\ s \end{pmatrix}_{L}, \quad c_{R}, \quad s_{R} \quad \begin{pmatrix} \nu_{\mu} \\ \mu^{-} \end{pmatrix}_{L}, \quad \mu_{R}^{-}$$

$$\begin{pmatrix} t \\ b \end{pmatrix}_{L}, \quad t_{R}, \quad b_{R} \quad \begin{pmatrix} \nu_{\tau} \\ \tau^{-} \end{pmatrix}_{L}, \quad \tau_{R}^{-}$$

$$3 generations of matter$$

Coupled through g, W^{\pm}, Z^0, γ

Gauge Sector

•Couplings of fermions to gauge bosons



determined by gauge symmetry

•Tested with sub-percent precision at LEP, Tevatron, SLD



Symmetry Breaking Sector

• M_Z, M_W, m_f require spontaneous symmetry breaking

$$(D_{\mu}\Phi)^{\dagger} D^{\mu}\Phi \Longrightarrow g^{2} \frac{v^{2}}{2} W_{\mu}^{+} W^{-\mu} + \frac{(g^{2} + g'^{2}) v^{2}}{4} Z_{\mu} Z^{\mu}$$

$$\text{when} \quad \langle \Phi \rangle = \begin{pmatrix} 0 \\ \frac{v}{\sqrt{2}} \end{pmatrix}$$

Also
$$Y_f \bar{f}_L \Phi f_R \implies Y_f \frac{v}{\sqrt{2}}$$

But SB sector much more open than gauge sector of the SM

6



Fit with $m_h = 126 \text{ GeV}$





Oblique Corrections



I.Where is the scalar sector coming from ?

•EWSB requires Higgs sector

•SM corresponds to minimal choice

 $\mathcal{L}_{\Phi} = \left(D_{\mu}\Phi\right)^{\dagger} D^{\mu}\Phi + V(\Phi^{\dagger}\Phi)$

with

$$V(\Phi^{\dagger}\Phi) = -m^2 \Phi^{\dagger}\Phi + \lambda (\Phi^{\dagger}\Phi)^2$$

Symmetry Breaking Sector



 $m_h = \sqrt{2\lambda}v$

Needs to be $m_h \leq O(1)$ TeV

to unitarize theory with $M_Z, M_W \simeq O(100) \text{ GeV}$





Where is the Scalar Sector Coming From

- •But what determines m and λ ?
- •Is the scalar sector resulting from some underlying dynamics ?
 - E.g. Superconductivity:
 - Cooper pairs $\Rightarrow \langle \Phi \rangle \neq 0$



The Higgs Mechanism and Superconductivity

Electromagnetism in a Superconductor

$$\mathcal{L} = -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} + (\partial_{\mu} + 2ieA_{\mu}) \Phi^{\dagger} (\partial_{\mu} - 2ieA_{\mu}) \Phi - V(\Phi^{\dagger}\Phi)$$

Complex scalar field Φ with U(1) gauge symmetry

$$\left. \begin{array}{c} \Phi \to e^{i\,\alpha(x)}\,\Phi \\ A_{\mu} \to A_{\mu} + \partial_{\mu}\alpha \end{array} \right\} \quad \Longrightarrow \quad \delta\mathcal{L} = 0$$

At $T < T_c$ $V(\Phi^{\dagger}\Phi) = -\mu^2 \Phi^{\dagger}\Phi + \lambda (\Phi^{\dagger}\Phi)^2$

The Higgs Mechanism and Superconductivity

Spontaneous breaking of U(1)

At $T < T_c$ $\langle \Phi \rangle \neq 0$ breaks EM

Photon acquires effective mass in the superconductor

In reality Φ is a condensate of electron pairs (Cooper)

 $\epsilon_{\alpha\beta} \Phi(x) = \langle 0 | \psi_{\alpha}(x) \, \psi_{\beta}(x) | 0 \rangle$

The Landau-Ginzburg description can be obtained from the microscopic theory of SC: BCS

II. Why is the Higgs so Light ?

 m_h not stable under radiative corrections



 $\Delta m_h^2 \simeq rac{c}{16\pi^2} \Lambda^2$ quadratically divergent

C determined by SM states: t, W^{\pm}, Z^0, h

The Hierarchy Problem

E.g. : The top quark contribution

Contributions from gauge bosons and h itself have similar form

The Hierarchy Problem

Renormalization group evolution of the Higgs mass

$$m_h^2(100 \text{ GeV}) = m_h^2(\Lambda) + \Delta m_h^2$$

UV physics =>
$$m_h(\Lambda)$$

 Λ
SM physics => $m_h(100 \text{ GeV})$
 100 GeV

But *c* is determined by SM fields at the EW scale

Need to adjust bare parameters (e.g. λ, m) to cancel these

$$\left(m_{h\text{bare}}^2 - \frac{c}{16\pi^2}\Lambda^2\right) \simeq O(100 \text{ GeV})^2$$

$$\implies$$
 Need fine tuning for $\Lambda \gg 1 \ TeV$

But physics that determines m_{hbare} lives above Λ

$$\Rightarrow$$
 Hierarchy Problem

III. Why are the fermion masses so different ?

Fermion masses come from Yukawa couplings

 $Y_f \, \bar{f}_L \Phi f_R$

But why $Y_t \simeq 1$, $Y_c \simeq 10^{-2}$, $Y_u \simeq 10^{-5}$ $Y_b \simeq 10^{-2}$, $Y_s \simeq 10^{-3}$, $Y_d \simeq 10^{-5}$ $Y_\tau \simeq 10^{-2}$, $Y_u \simeq 10^{-3}$, $Y_e \simeq 10^{-6}$

IV. How do neutrinos get masses ?

If we want m_{ν} from $Y_{\nu} \bar{\nu}_L \Phi \nu_R$ we need ν_R

But
$$Y_{\nu_R} = 0, \ T^3_{\nu_R} = 0, \ Q(\nu_R) = 0$$

 $\Rightarrow \nu_R$ has no SM interactions !

The Problem(s) with the Standard Model IV. How do neutrinos get masses ? (cont.) If neutrinos are Majorana particles we can have $-m_{\nu} \bar{\nu}_{L}^{c} \nu_{L} + \text{h.c.}$

From the dim-5 operator $\frac{k}{\Lambda} \bar{L} \tilde{\Phi} \tilde{\Phi} L$

leading to $\frac{k}{\Lambda}$

$$\frac{k}{\Lambda} v^2 \ \bar{\nu}_L^c \ \nu_L$$

New physics scale $\Lambda \simeq O(10^{15}) \text{ GeV}$ to get m_{ν}

III. Dark Matter

We know that $\begin{array}{l} \Omega_{\Lambda} \simeq 0.73 \\ \Omega_{CDM} \simeq 0.23 \\ \Omega_{b} \simeq 0.04 \end{array}$

- Most of matter is non-baryonic cold dark matter
- The SM does not have a suitable DM candidate
- •Need new physics beyond the SM to explain CDM

Dark Matter



- Neutrinos: too light and hot
- Axions: very light $(m_a \simeq 10^{-5} {\rm eV})$, very little interaction
- Weakly Interacting Massive Particles (WIMPs):

 $m_{\chi} \simeq (1 - 1000) \text{GeV}$

Dark Matter

WIMP Coincidence

• If WIMPs are thermal relics

 $\chi + \bar{\chi} \leftrightarrow \mathrm{SM} \quad \text{for} \quad \Gamma \gg H$

• Freeze out:

for $\ \Gamma < H$ annihilation of $\chi_{'{
m s}}$ stops

$$n_{\chi} \simeq a^{-3}$$

• WIMP Relic Density:

$$\Omega_{\chi} h^2 = \frac{m_{\chi} n_{\chi}}{\rho_c} \simeq 3 \times 10^{-27} \text{cm}^3 \text{ s}^{-1} \frac{1}{\langle \sigma_A v \rangle}$$

Dark Matter

WIMP Relic Density:

$$\Omega_{\chi} h^2 = \frac{m_{\chi} n_{\chi}}{\rho_c} \simeq 0.1 \text{pb} \cdot \text{c} \frac{1}{\langle \sigma_A v \rangle}$$

For a typical weakly interacting particle

$$\langle \sigma_A v \rangle \simeq \frac{\alpha^2}{m_\chi^2} \ c \ \simeq 1 \, \mathrm{pb} \cdot \mathrm{c} \quad \mathrm{for} \quad m_\chi \sim 100 \ \mathrm{GeV}$$



Other Problems

What's the origin of the baryon asymmetry ?

The Strong CP Problem



Not necessarily associated with the Symmetry Breaking Sector

What Physics Beyond the Standard Model

• <u>Solves a problem:</u>

Origin of the scalar sector Gauge hierarchy problem Fermion mass hierarchy

• Experimentally accessible

We'll see it at the LHC or close

Physics Beyond the Standard Model

Organize by origin of Higgs sector or solution to HP

•<u>Supersymmetry:</u>

Higgs is elementary SUSY protects m_h

• <u>Higgs sector is composite:</u>

Technicolor. No Higgs. X

Higgs is a pNGB

Physics Beyond the Standard Model

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Beyond the Standard Model II - SUSY

- •Supersymmetry: a solution to the Hierarchy Proble
- •Basic elements of SUSY theories
 - •The MSSM
- •The MSSM and the Higgs

Supersymmetry and the Hierarchy Problem

Protecting Fermion Masses: Chiral Symmetry

Fermion masses only log divergent. E.g. QED



Chiral symmetry protects m_e to all orders in PT

I.
$$\delta m_e \longrightarrow 0$$
 for m_e^0

2. Divergence is logarithmic

Supersymmetry and the Hierarchy Problem

How to protect the Higgs mass ?

Introduce a fermionic partner of the Higgs: Higgsino

Need symmetry to relate Higgs (boson) to Higgsino (fermion)

 \Rightarrow Supersymmetry

Higgs and Higgsino form a SUSY multiplet (H, \tilde{H})



Supersymmetry and the Hierarchy Problem What about the top quark Λ^2 divergence ? All fermions will have a scalar partner and viceversa stop quark \tilde{t} forms SUSY multiplet with t (t, \tilde{t})



No divergences in exact SUSY

Supersymmetric Theories

Matter in Chiral Supermultiplets: Complex scalar Weyl fermion

Gauge in Vector Supermultiplets: Weyl fermion
Supersymmetric Theories

SUSY transformations turn scalars into fermions and viceversa leaving the lagrangian invariant

$$\mathcal{L} = \partial_{\mu}\phi^*\partial^{\mu}\phi + i\psi^{\dagger}\bar{\sigma}_{\mu}\partial^{\mu}\psi$$

 $\sigma_0 = \bar{\sigma}_0 = I$ $\sigma_i = -\bar{\sigma}_i$

 $\implies \qquad \qquad \delta\phi = \epsilon \, \psi$

with $\epsilon\,$ fermionic anti-commuting infinitesimal change and $\delta\psi=i\sigma^\mu\epsilon\partial_\mu\phi$

Supersymmetric Theories

<u>Superspace</u>

Coordinates $y^{\mu} = x^{\mu} - \theta \bar{\sigma}^{\mu} \bar{\theta}$

 θ : two-component Grassmann spinor $\theta_{\alpha}, \quad \theta_{\alpha}^{\dagger} \equiv \bar{\theta}_{\dot{\alpha}}$ Chiral superfield

$$\begin{split} \Phi(y) &= \phi(y) + \sqrt{2} \,\theta \psi(y) + \theta^2 \,F(y) \\ &= \phi(x) - i\theta \sigma^\mu \bar{\theta} \,\partial_\mu \phi(x) - \frac{1}{4} \,\theta^2 \bar{\theta}^2 \,\partial^2 \phi(x) \\ &+ \sqrt{2}\theta \,\psi(x) + \frac{i}{\sqrt{2}} \,\theta^2 \partial_\mu \psi(x) \sigma^\mu \bar{\theta} + \theta^2 \,F(x) \end{split}$$

SUSY in Superspace

- θ and $\overline{\theta} \Rightarrow \theta^n = 0$ for $n \ge 3$
- $\int d^2\theta \,\theta^2 = 1$ selects coefficient of θ^2

•
$$d^4\theta \equiv d^2\theta \, d^2\bar{\theta} \implies \int d^4\theta$$
 selects coefficient of $\theta^2 \, \bar{\theta}^2$

•The θ^2 component of a CSF is a total derivative under SUSY $\Rightarrow \int d^2 \theta W(\Phi)$ is SUSY invariant

•Same for $\theta^2 \bar{\theta}^2$ components $\Rightarrow \int d^4 \theta K(\Phi^{\dagger}, \Phi)$ invariant under SUSY

SUSY in Superspace

E.g. Kinetic terms in free theory

$$\int d^4\theta \, \Phi^{\dagger} \Phi = \partial_{\mu} \phi^* \partial^{\mu} \phi + i \psi^{\dagger} \bar{\sigma}^{\mu} \partial_{\mu} \psi + F^* F + \text{total derivatives}$$
$$= \mathcal{L}_{\text{free}}$$

Superpotential $W(\Phi)$: Generates interactions through

$$\int d^2\theta \, W(\Phi) = \mathcal{L}_{\rm int}$$

where $W(\Phi)$ is holomorphic function of Φ

SUSY in Superspace

Gauge Superfields

$$V^{a}_{\mu} = \theta \bar{\sigma}^{\mu} \bar{\theta} A^{a}_{\mu} + i \theta^{2} \bar{\theta} \lambda^{a\dagger} - i \theta \bar{\theta}^{2} \lambda^{a} + \frac{\theta^{2} \bar{\theta}^{2}}{2} D^{a}$$

Gauge transformation for gauge superfields

 $e^{t^a V^a} \rightarrow e^{t^a \Lambda^{a\dagger}} e^{t^a V^a} e^{t^a \Lambda^a}$ Λ^a : gauge parameter is superfield

$$\Rightarrow V^a \to V^a + \Lambda^{a\dagger} + \Lambda^a + O(V^a \Lambda^a)$$

For chiral superfields:

$$\Phi \to e^{-gt^a \Lambda^a} \Phi$$

SUSY Interactions

Gauge-invariant kinetic terms

$$\int d^{4}\theta \, \Phi^{\dagger} \, e^{gt^{a}V^{a}} \, \Phi = (D_{\mu}\phi)^{\dagger}D^{\mu}\phi + i\psi^{\dagger}\bar{\sigma}^{\mu} \, D_{\mu}\psi$$
$$-\sqrt{2}g \left[(\phi^{*}t^{a}\psi) \,\lambda^{a} + \lambda^{a\dagger}(\psi^{\dagger}t^{a}\phi) \right]$$
$$+g(\phi^{*}t^{a}\phi)D^{a}$$

In addition to usual gauge interactions



SUSY Interactions

Gauge fields kinetic terms: superfield strength

$$\mathcal{W}^{a} = -\sigma^{\mu\nu}\theta F^{a}_{\mu\nu}(y) - \theta^{2}\sigma_{\mu}D^{\mu}\lambda^{a}(y) - i\lambda^{a}(y) + \theta D^{a}(y)$$

is a chiral superfield



Supersymmetric Theories

<u>Summary</u>

•Gauge and SUSY invariant kinetic terms for matter

$$\int d^4\theta \, \Phi^\dagger \, e^{gt^a V^a} \, \Phi$$

•Gauge and SUSY invariant kinetic terms for gauge fields

$$\int d^2\theta \, \mathcal{W}^a(y) \, \mathcal{W}^a(y)$$

•Gauge and SUSY invariant non-gauge interactions $\int d^2\theta \, W(\Phi)$

Supersymmetry

Supersymmetric extension of the SM





Supersymmetry

<u>MSSM</u>

•Interactions still determined by SM gauge $SU(3)_C imes SU(2)_L imes U(1)_Y$



Supersymmetry

•<u>Superpotential</u>

$W_{\rm MSSM} = \bar{u}Y_uQH_u - \bar{d}Y_dQH_d - \bar{e}Y_eLH_d + \mu H_uH_d$



Soft SUSY Breaking

• Need to break SUSY softly:

$$\begin{split} W_{\text{soft}} &= -\frac{1}{2} \left(M_1 \tilde{B} \tilde{B} + M_2 \tilde{W} \tilde{W} + M_3 \tilde{g} \tilde{g} + \text{h.c.} \right) \\ &- \tilde{Q}^{\dagger} \, m_Q^2 \, \tilde{Q} - \tilde{L}^{\dagger} \, m_L^2 \, \tilde{L} - \tilde{\bar{u}} \, m_u^2 \, \tilde{\bar{u}}^{\dagger} - \tilde{\bar{d}} \, m_d^2 \, \tilde{\bar{d}}^{\dagger} - \tilde{\bar{e}} \, m_e^2 \, \tilde{\bar{e}}^{\dagger} \\ &- \left(\tilde{\bar{u}} \, A_u \, \tilde{Q} H_u - \tilde{\bar{d}} \, A_d \, \tilde{Q} H_d + \tilde{\bar{e}} \, A_e \, \tilde{L} H_d + \text{h.c.} \right) \\ &- m_{H_u}^2 H_u^* H_u - m_{H_d}^2 H_d^* H_d - (b H_u H_d + \text{h.c.}) \end{split}$$

R Parity

Additional SUSY-preserving terms in the superpotential

 $W_{\rm RPV} = \alpha^{ijk} Q_i L_j \bar{d}_k + \beta^{ijk} L_i L_j \bar{e}_k + \gamma^i L_i H_u + \delta^{ijk} \bar{d}_i \bar{d}_j \bar{u}_k$

they violate B and L !



R Parity

Introduce new discrete symmetry, M parity

 $P_M = (-1)^{3(B-L)}$

Forbids terms W that violate B, L

• Equivalent to R parity

$$P_R = (-1)^{3(B-L)+2s}$$

$$\Rightarrow \begin{cases} Superpartners have $P_R = -1 \\ SM \text{ particles have } P_R = +1 \end{cases}$$$

R Parity

Lightest Supersymmetric Particle (LSP) is stable



decay of LSP forbidden by R parity

Typical SUSY WIMP candidate: neutralino: $\tilde{\chi}^0$ admixture of $\tilde{W}, \ \tilde{B}, \ \tilde{H}$

In generic SUSY models is possible to obtain the correct Ω_{χ}

Implications of m_h for SUSY



Superpartner loops cancel quadratic divergences

$$m_h^2 = m_Z^2 \,\cos^2 2\beta + \frac{3m_t^4}{4\pi v^2} \left(\log\left(\frac{M_S^2}{m_t^2}\right) + \frac{X_t^2}{M_S^2} \,\left(1 - \frac{X_t^2}{M_S^2}\right) \right)$$

 $M_S = \sqrt{m_{\tilde{t}_1} m_{\tilde{t}_2}}$ Stop mass scale

 $X_t = A_t - \mu \cot \beta \qquad \text{Stop mixing}$

SUSY Phenomenology

MSSM with R parity conservation

• E.g. $pp \to \tilde{g}\tilde{g}, \ \tilde{q}\tilde{q}^*, \ \tilde{q}\tilde{q}$ with $\tilde{q} \to q \chi_1^0$ or $\tilde{g} \to \tilde{q} q$

 \implies jets + $E_T^{\text{miss.}}$

- Or 3-body decays. E.g. $\tilde{g} \rightarrow q \, \bar{q} \, \chi_1^0$
- •Also decays with I or more leptons
- •Bounds depend on decay channels/models





•Assuming direct decays to jets



• Assume $\tilde{g} \rightarrow q \, \bar{q} \, \chi_1^0$



				<i>,</i>
	MSUGBA/CMSSM : 0 len + i's + F-	/ -5.8 fb ⁻¹ .8 TeV [ATLAS_CONE_2012_100]		
θS	MSUGRA/CMSSM : 1 len + i's + F	$L = 5.0 \text{ H}^{-1}$ 8 ToV [ATLAS CONE 2012 103]		
rch	Pheno model : 0 len $\pm i$'s $\pm F$	L=5.0 10 , 0 10V [ATLAS-CONF-2012-104]	$Ldt : C_{1,24} = C_{$	= (1.00 - 5.8) fb ⁻ '
eai	Phono model : 0 lop $\pm 13 \pm L_{T,\text{miss}}$	L=5.6 ID , 6 IEV [ATLAS-CONF-2012-109]	1.16 TeV g mass $(m(q) < 2 \text{ TeV}, \text{ light}_{1})$	
S C	Cluips mod $\tilde{z}^{\pm} (\tilde{z} + s \bar{z}^{-1}) + 1$ log $z = 1$	L=5.8 fb , 8 lev [AILAS-CONF-2012-109]	1.38 lev q mass $(m(g) < 2 \text{ lev}, \text{ light } \chi)$	13 - 7,0100
sive	Giulito med. χ (g \rightarrow qq χ): Tiep + JS + E _{T,miss}	L=4.7 fb , 7 lev [AILAS-CONF-2012-041]	900 GeV 9 IIIASS $(m(\chi_1) < 200 \text{ GeV}, m(\chi_1) = \frac{1}{2}(m(\chi_1) + m(g))$	ATIAS
sinic	GIVISB: 2 IEP (US) + JS + $E_{T,miss}$	L=4.7 fb ⁻ , 7 TeV [Preliminary]	1.24 TeV G ($\tan\beta < 15$)	Broliminory
lnc	GIVISD: $1-2^{1} + 0 - 1$ lep + $js + E$	L=4.7 fb ⁻¹ , 7 TeV [ATLAS-CONF-2012-112]	1.20 TeV g mass $(\tan\beta > 20)$	Freinninary
	$\alpha = 0$	L=4.8 fb ⁻¹ , 7 TeV [ATLAS-CONF-2012-072]	1.07 TeV g mass $(m(\tilde{\chi}_1) > 50 \text{ GeV})$	
	$\tilde{g} \rightarrow bb \tilde{\chi}_{1}$ (virtual b) : 0 lep + 1/2 b-j's + $E_{T,miss}$	L=2.1 fb ⁻¹ , 7 TeV [1203.6193]	900 GeV g mass $(m(\chi_1) < 300 \text{ GeV})$	
d (S	$\widetilde{g} \rightarrow b\overline{b}\widetilde{\chi}_{\perp}^{\circ}$ (virtual b) : 0 lep + 3 b-j's + $E_{T,miss}$	L=4.7 fb ⁻¹ , 7 TeV [1207.4686]	1.02 TeV \widetilde{g} mass $(m(\widetilde{\chi}_{d}) < 400 \text{ GeV})$	
arl	$\tilde{g} \rightarrow b \tilde{b} \tilde{\chi}_{1}^{\circ}$ (real b) : 0 lep + 3 b-j's + $E_{T,miss}$	L=4.7 fb ⁻¹ , 7 TeV [1207.4686]	1.00 TeV \tilde{g} mass $(m(\tilde{\chi}_1^0) = 60 \text{ GeV})$	
squ	$\widetilde{g} \rightarrow t t \widetilde{\chi}_{10}^{0}$ (virtual t) : 1 lep + 1/2 b-j's + $E_{T,miss}$	L=2.1 fb ⁻¹ , 7 TeV [1203.6193]	710 GeV $\widetilde{\mathbf{g}}$ mass $(m(\widetilde{\chi}_1^0) < 150 \text{ GeV})$	
7. S Me	$\tilde{g} \rightarrow t \tilde{\chi}^{0}$ (virtual \tilde{t}) : 2 lep (SS) + j's + $E_{T \text{ miss}}$	L=5.8 fb ⁻¹ , 8 TeV [ATLAS-CONF-2012-105]	850 GeV \widetilde{g} mass $(m(\widetilde{\chi}_1^0) < 300 \text{ GeV})$	
gei no	$\widetilde{g} \rightarrow t \widetilde{\chi}^0$ (virtual \widetilde{t}) : 3 lep + j's + $E_{\tau \text{ miss}}$	L=4.7 fb ⁻¹ , 7 TeV [ATLAS-CONF-2012-108]	760 GeV \widetilde{g} mass (any $m(\widetilde{\chi}_{i}^{0}) < m(\widetilde{g})$)	
rd	$\widetilde{g} \rightarrow tt \widetilde{\chi}^{0}$ (virtual \widetilde{t}) : 0 lep + multi-j's + $E_{\tau \text{ miss}}$	L=5.8 fb ⁻¹ , 8 TeV [ATLAS-CONF-2012-103]	1.00 TeV g̃ mass (m(χ̃) < 300 GeV)	
6 0	$\widetilde{q} \rightarrow t\widetilde{t} \widetilde{\chi}^0$ (virtual \widetilde{t}) : 0 lep + 3 b-i's + $E_{T,miss}$	L=4.7 fb ⁻¹ , 7 TeV [1207.4686]	940 GeV \widetilde{g} mass $(m(\widetilde{\chi}^0)^2 < 50 \text{ GeV})$	
	$\widetilde{q} \rightarrow t\widetilde{t} \widetilde{\gamma}^{0}$ (real \widetilde{t}) : 0 lep + 3 b-i's + $E_{T,miss}$	L=4.7 fb ⁻¹ , 7 TeV [1207.4686]	820 GeV \widetilde{g} mass $(m(\widetilde{\chi}^0) = 60 \text{ GeV})$	
	$\widetilde{bb}, \widetilde{b} \rightarrow \widetilde{b\gamma}^{0}$: 0 lep + 2-b-iets + E_{τ}	L=4.7 fb ⁻¹ , 7 TeV [ATLAS-CONF-2012-106]	480 GeV b mass $(m(\tilde{\chi}^0) < 150 \text{ GeV})$	
ks nn	$\widetilde{bb}, \widetilde{b} \rightarrow t\widetilde{\gamma}^{\pm}$; 3 lep + i's + E_{\pm}	L=4.7 fb ⁻¹ . 7 TeV [ATLAS-CONF-2012-108]	380 GeV $\widetilde{\mathbf{q}}$ mass $(m(\widetilde{\chi}^{\pm}) = 2m(\widetilde{\chi}^{0}))$	
ctic	\widetilde{t} (very light), $\widetilde{t} \rightarrow b \widetilde{v}^{\pm}$; 2 lep + E_{-}	L=4.7 fb ⁻¹ , 7 TeV [CONF-2012-059]135 GeV	\tilde{t} mass $(m\tilde{\chi}^0) = 45 \text{ GeV}$	
npa 1bs	$\widetilde{\mathrm{ff}}$ (light) $\widetilde{\mathrm{f}} \rightarrow \mathrm{b}\widetilde{\mathrm{v}}^{\pm} \cdot 1/2$ len + h-iet + F	$l = 4.7 \text{ fb}^{-1}$, 7 TeV [CONF-2012-070] 120-173 G	\tilde{t} mass $(m(\tilde{\chi}^0) = 45 \text{ GeV})$	
n. , pro	$\widetilde{T}_{T,\text{miss}}$	$I = 4.7 \text{ fb}^{-1}$ 7 TeV [1208 1447]	380-465 Cav T mass $(m(\tilde{\omega}^0) = 0)$	
ge	\widetilde{tt} (heavy), $\widetilde{t} \rightarrow \widetilde{tx}_{0}^{\circ}$: 1 lep + b-jet + E	$l = 4.7 \text{ fb}^{-1}$ 7 TeV [CONE-2012-073]	230-440 GeV $t \text{ mass } (m(x_1^0) = 0)$	
3rd dire	\widetilde{tt} (heavy), \widetilde{t} , \widetilde{tt} ? len + b jet + E	$l = 4.7 \text{ fb}^{-1}$ 7 TeV [CONE-2012-070]	200-305 GoV t mass $(mc_1^{(0)}) = 0$	
0 0	$\widetilde{H} (GMSR)^{1} Z (\rightarrow II) + h_{i} o f I Z (\rightarrow II) I I I I I I I I$	$L = 9.1 \text{ fb}^{-1}$ 7 TeV [1004 6726]	230-303 GeV T mass $(m(\chi_1) = 0)$	
	$T_{1} = T_{1} = T_{1}$	L=2.110, 7 TeV [1204.0730]	$mass (mC^0) = 0$	
W ect	$\Pi_{L}^{1}, \Pi \rightarrow I\chi : 2 \text{ lep } + \mathbb{Z}_{T,\text{miss}}$	L=4.7 fb , 7 lev [CONF-2012-076] 93-180 ($\frac{1}{10000000000000000000000000000000000$	
Ч	$\chi_1 \chi_1, \chi_1 \rightarrow \text{IV}(\text{IV}) \rightarrow \text{IV}\chi_1$. 2 lep + $E_{T,\text{miss}}$	L=4.7 fb , 7 lev [CONF-2012-076]	120-330 GeV χ_1 midds $(m(\chi_1) = 0, m(l,v) = (m(\chi_1) + m(\chi_1)))$	
	$\chi \chi \rightarrow 31(Ivv) + v + 2\chi_1)$ 3 lep + $E_{T, miss}$	L=4.7 fb , 7 lev [CONF-2012-077]	60-500 GeV χ_1 mass $(m(\chi_1) = m(\chi_2), m(\chi_1) = 0, m(l,v)$ as above)	
p∈s	AMSB (direct χ_1 pair prod.) : long-lived χ_1	L=4.7 fb ⁻¹ , 7 TeV [ATLAS-CONF-2012-111] 21	10 GeV χ_1 IIIaSS $(1 < \tau(\chi_1) < 10 \text{ ns})$	
live cle	Stable g R-hadrons : Full detector	L=4.7 fb ⁻¹ , 7 TeV [ATLAS-CONF-2012-075]	985 GeV g mass	
ng- artii	Stable t R-hadrons : Full detector	L=4.7 fb ⁻¹ , 7 TeV [ATLAS-CONF-2012-075]	683 GeV t mass	
Lo. Dê	Metastable g R-hadrons : Pixel det. only	L=4.7 fb ⁻¹ , 7 TeV [ATLAS-CONF-2012-075]	910 GeV g mass $(\tau(g) > 10 \text{ ns})$	
	GMSB : stable τ	L=4.7 fb ⁻¹ , 7 TeV [ATLAS-CONF-2012-075]	310 GeV τ Mass (5 < tan β < 20)	
	RPV : high-mass eµ	L=1.1 fb ⁻¹ , 7 TeV [1109.3089]	1.32 TeV v_{τ} MASS $(\lambda_{311}^{*}=0.10, \lambda_{312}^{*}=0.05)$	
_ ∠	Bilinear RPV : 1 lep + j's + $E_{T,miss}$	L=1.0 fb ⁻¹ , 7 TeV [1109.6606]	760 GeV $\tilde{\mathbf{q}} = \tilde{\mathbf{g}} \text{ mass} (c\tau_{LSP} < 15 \text{ mm})$	
6	BC1 RPV : 4 lep + $E_{T,miss}$	L=2.1 fb ⁻¹ , 7 TeV [ATLAS-CONF-2012-035]	1.77 TeV g mass	
	RPV $\widetilde{\chi}_{1}^{\upsilon} \rightarrow qq\mu : \mu + heavy displaced vertex$	L=4.4 fb ⁻¹ , 7 TeV [ATLAS-CONF-2012-113]	700 GeV $\tilde{\mathbf{q}}$ mass $(3.0 \times 10^{-6} < \lambda_{211} < 1.5 \times 10^{-5}, 1 \text{ mm} < c\tau < 1 \text{ m}, \tilde{\mathbf{g}})$	decoupled)
<u>e</u>	Hypercolour scalar gluons : 4 jets, $m_{ij} \approx m_{kl}$	L=4.6 fb ⁻¹ , 7 TeV [ATLAS-CONF-2012-110]	100-287 GeV Sgluon mass (incl. limit from 1110.2693)	
)the	Spin dep. WIMP interaction : monojet + $\tilde{E}_{T,miss}$	L=4.7 fb ⁻¹ , 7 TeV [ATLAS-CONF-2012-084]	709 GeV M [*] SCale (m_{χ} < 100 GeV, vector D5, Dirac χ)	
^O S	pin indep. WIMP interaction : monojet $+E_{T,miss}$	L=4.7 fb ⁻¹ , 7 TeV [ATLAS-CONF-2012-084]	548 GeV M^* SCale $(m_{\chi} < 100 \text{ GeV}, \text{ tensor D9}, \text{Dirac}\chi)$	
	5 N Y Y Y			
		10 ⁻¹	1 10	
		10	i 10	
*Only a selection of the available mass limits on new states or phenomena shown. Mass scale [TeV				

ATLAS SUSY Searches* - 95% CL Lower Limits (Status: SUSY 2012)

*Only a selection of the available mass limits on new states or phenomena shown. All limits quoted are observed minus 1σ theoretical signal cross section uncertainty.

Hiding SUSY

Why haven't we seen it ?

•Compressed Spectrum

Not enough $E_T^{\text{miss.}}$

•R-parity Violation

LSP not stable. Different decay modes. Not enough $E_T^{\text{miss.}}$

•Natural SUSY

Light higgsinos, 3rd. gen. squarks Everybody else heavy

Natural SUSY

Naturalness only requires Higgsinos, stops and gluinos to be "light"



Natural SUSY

It's hard to produce light stops



Hiding SUSY

Stop limits



Finding Natural SUSY is hard

Implications of m_h for SUSY



Superpartner loops to make Higgs heavier

$$m_h^2 = m_Z^2 \, \cos^2 2\beta + \frac{3m_t^4}{4\pi v^2} \left(\log\left(\frac{M_S^2}{m_t^2}\right) + \frac{X_t^2}{M_S^2} \, \left(1 - \frac{X_t^2}{M_S^2}\right) \right)$$

 $M_S = \sqrt{m_{\tilde{t}_1} m_{\tilde{t}_2}}$ Stop mass scale

 $X_t = A_t - \mu \cot \beta \qquad \text{Stop mixing}$

SUSY and the Higgs

$$m_h^2 = m_Z^2 \, \cos^2 2\beta + \frac{3m_t^4}{4\pi v^2} \left(\log\left(\frac{M_S^2}{m_t^2}\right) + \frac{X_t^2}{M_S^2} \, \left(1 - \frac{X_t^2}{M_S^2}\right) \right)$$

For $m_h = 125 \text{ GeV}$

 $\Rightarrow \tan \beta > 3.5$



SUSY and the Higgs



to get large enough superpartner masses

Beyond the MSSM

Problem in the MSSM:

$$V(H_u, H_d) = \frac{(g^2 + g'^2)}{2} (H_u^2 - H_d^2)^2 \qquad \Rightarrow m_h^2 = M_Z^2 \cos^2(2\beta)$$

$\underline{\mathsf{NMSSM}} \quad \mathsf{Add} \text{ a singlet chiral superfield} \\ \lambda_S \, S \, H_u \, H_d$

$$\langle S \rangle = v_s \quad \Rightarrow \quad \lambda_S v_s H_u H_d \qquad \text{gives } \mu \text{ term}$$

and an extra quartic $\lambda_S^2 H_u^2 H_d^2$

$$\implies m_h^2 = M_Z^2 \cos^2(2\beta) + \lambda_S^2 v^2 \sin^2(2\beta) + \cdots$$

SUSY - Conclusions/Outlook

- •SUSY is a beautiful solution to the Hierarchy Problem
- •The MSSM spectrum is highly constrained if we want ${ ilde m}_Q \leq O(1)~{
 m TeV}$
- But natural spectrum very much viable
- Bottom-up approach: look for natural SUSY signals if we really want to exclude SUSY
 - •The measurement of m_h posses additional constraints.
 - •Extensions of the MSSM (NMSSM, extended gauge sectors) should be explored, as long as they remain natural solutions to the HP

Physics Beyond the Standard Model 3.1

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Arequipa, Peru, March 6-19 2013

Beyond the Standard Model

Lecture I •Why do we need to go Beyond the SM ? •The Hierarchy Problem: what do we need to solve it ?

Lecture 2 • Supersymmetry and the Hierarchy Problem

Lecture 3 •New Dynamics at the TeV scale: the Higgs as a (pseudo) Nambu-Goldstone Boson

Beyond the Standard Model III

- Solve the Hierarchy problem with dynamics: QCD and the σ (Technicolor, ...)
- Dynamical (composite) <u>light</u> Higgs: is a (pseudo) Goldstone boson The example of the pion in QCD
 - •Composite Higgs Models:

Little Higgs Twin Higgs Gauge-Higgs unification in AdS_5 Where is the Scalar Sector Coming From

- •But what determines m and λ ?
- •Is the scalar sector resulting from some underlying dynamics ?
 - E.g. Superconductivity:
 - Cooper pairs $\Rightarrow \langle \Phi \rangle \neq 0$



Physics Beyond the Standard Model

Organize by origin of Higgs sector or solution to HP

•<u>Supersymmetry:</u>

Higgs is elementary SUSY protects m_h

• <u>Higgs sector is composite:</u>

Technicolor. No Higgs. X

Higgs is a pNGB
Composite Scalars: the Example of QCD Spontaneous breaking of chiral symmetry in QCD

QCD with 2 flavors:

$$\mathcal{L}_{\text{QCD}} = \bar{Q}_L \, i \, \not \!\!\! D Q_L + \bar{Q}_R \, i \, \not \!\!\! D Q_R - \bar{Q}_L \, M Q_R + \text{h.c.}$$

with

$$Q = \begin{pmatrix} u \\ d \end{pmatrix} \qquad \qquad M = \begin{pmatrix} m_u & 0 \\ 0 & m_d \end{pmatrix}$$

If M = 0, is invariant under $SU(2)_L \times SU(2)_R$

$$\begin{array}{ll} Q_{L} & \longrightarrow & e^{i\ell^{a}t^{a}} Q_{L} \\ Q_{R} & \longrightarrow & e^{ir^{a}t^{a}} Q_{R} \end{array} & \text{with} & \begin{cases} t^{a} = \frac{\sigma^{a}}{2}, & a = 1, 2, 3 \\ \ell^{a}, r^{a} & \text{free parameters} \end{cases} \end{array}$$

Chiral Symmetry Breaking



At low energies, $\Lambda \simeq \Lambda_{\rm QCD}$, quark condensation

 $\langle \bar{Q}_L Q_R \rangle \neq 0 \implies SU(2)_L \times SU(2)_R \longrightarrow SU(2)_V$

• Quarks acquire a dynamical mass

 $m_Q \sim \Lambda_{\rm QCD}$

Chiral Symmetry Breaking

•3 broken generators \implies 3 NGBs (π^+, π^-, π^0)

Since $SU(2)_L \times SU(2)_R = SU(2)_V \times SU(2)_A \longrightarrow SU(2)_V$

Axial current $j^{a5}_{\mu} = \bar{Q}\gamma_{\mu}\gamma^{5}Q$

does not annihilate the vacuum

$$\langle 0|j^{a5}_{\mu}|\pi^{b}(p_{\mu})\rangle = if_{\pi}\,p_{\mu}\,\delta^{ab}$$

But still conserved if $m_{\pi} = 0$

$$\partial^{\mu} j^{a5}_{\mu} = f_{\pi} m_{\pi}^2$$

Spontaneous Breaking of Chiral Symmetry

Linear σ model

$$\mathcal{L} = \frac{1}{4} \operatorname{Tr} \left[\partial_{\mu} \Sigma^{\dagger} \partial^{\mu} \Sigma \right] + \frac{\mu^2}{4} \operatorname{Tr} \left[\Sigma^{\dagger} \Sigma \right] - \frac{\lambda}{16} \left(\operatorname{Tr} \left[\Sigma^{\dagger} \Sigma \right] \right)^2$$

with $\Sigma = \sigma + i t^a \pi^a$ and $\Sigma \to L^\dagger \Sigma R$

If
$$\mu^2 > 0 \implies \langle \Sigma \rangle = v \neq 0$$
 $v = \sqrt{\frac{\mu^2}{\lambda}}$



Spontaneous Breaking of Chiral Symmetry

In real QCD:

• $\begin{cases} m_{\sigma} \sim \Gamma_{\sigma} \sim O(1) \text{ GeV } \text{Cutoff of the effective theory} \\ \sigma \text{ is not a low energy state (too broad to be observable)} \end{cases}$

- $m_u, m_d \neq 0 \implies$ Explicit symmetry breaking
 - $m_{\pi} \neq 0$ $\pi's$ are pseudo NGBs

But still light

$$m_{\pi} \simeq 0.14 \text{GeV} \ll O(1) \text{ GeV}$$

GeV vs. TeV Scales

Build a TeV-scale model of EWSB in analogy with QCD

Two avenues:

- Fermionic sector breaks EWS just as in QCD Higgs (σ) is not is the light spectrum
- Strong sector breaks global symmetry Higgs is a (pseudo) NGB remnant just like the $\pi's$

Strong Dynamics at the TeV Scale

Scaled up QCD

- •New gauge interaction
- •Strong at the TeV scale
- •Breaks EWS by

 $\langle \bar{F}F \rangle \neq 0$



Basic Technicolor Model

•Asymptotically-free interaction $SU(N_T)$

•New fermions: $SU(2)_L$ doublet

$Q_L = \left(\begin{array}{c} T\\ B \end{array}\right)_L$	$\left(N_T,1,2,Y_Q ight)$
T_R	$\left(N_T,1,1,Y_T ight)$
B_R	$\left(N_{T},1,1,Y_{B} ight)$

•At Λ_{TC} we have $\langle \bar{Q}_L Q_R \rangle \neq 0$



Higgs Mechanism without a Higgs

 $SU(2)_L \times SU(2)_R \rightarrow SU(2)_V \implies$ 3 Nambu-Goldstone bosons

NGBs eaten as gauge boson longitudinal polarizations

$$W_{\mu} \qquad W_{\nu} \qquad W_{\mu} \qquad W_{\nu} \qquad W_{\nu$$

$$i\frac{g^2 F_T^2}{4} \left(g_{\mu\nu} - \frac{q_\mu q_\nu}{q^2}\right)$$

Fermion Masses without a Higgs

Need extended interaction mixing SM fermions with tfermions



Extended Technicolor

ETC requires more techni-fermions

- $\begin{pmatrix} T \\ B \end{pmatrix}_{L}^{i} \qquad T_{R}^{i}, B_{R}^{i} \qquad \text{techni-quarks}$ $\begin{pmatrix} N \\ E \end{pmatrix}_{L} \qquad N_{R}, E_{R} \qquad \text{techni-leptons}$
 - Number of doublets higher $N_D = 4$ Problems with EWPC
 - Larger chiral symmetry broken $SU(8)_L \times SU(8)_R \longrightarrow SU(8)_V$ 63 -3 = 60 NGBs left in the spectrum!

Flavor Violation from ETC Interactions

ETC leads to tree-level flavor violation



$$\Rightarrow$$
 effects in $(K^0 - \overline{K}^0)$, $(B^0 - \overline{B}^0)$, mixing,...

 $\implies M_{\rm ETC} > 1000 {
m TeV}$

But $M_{\rm ETC}$ cannot be too large or it would suppress m_t, m_b, m_c too much

Walking Technicolor and Separation of Scales

To get heavier masses need to enhance TC condensate

 $\Rightarrow \begin{cases} \text{Near-conformal behavior of TC interaction} \\ \text{Coupling walks} \end{cases}$

But walking takes long time for coupling to become super-critical

 \implies Walking generates large separation of scale

Electroweak Precision Constraints

For the simple scaled up QCD scenario



New Ideas in Techni-Color Models

• Minimal Walking Technicolor (F. Saninno et al.)

 $N_T = 2, N_D = 1$

No flavor theory

Not clear how to get a light Higgs

Can be modeled in AdS_5

• Conformal Technicolor (M. Luty *et al.*) Strong sector is near a conformal fixed point in the UV Explicit conformal breaking \rightarrow EWSB First basic models accommodate light Higgs as pNGB

Higgs is a pseudo Nambu-Goldstone Boson Back to the analogy of QCD at low energies

• Build models where the Higgs is like π instead of σ Need to break global symmetry spontaneously $SU(2)_L \times SU(2)_R \rightarrow SU(2)_V$

Number of NGBs: 3 + 3 - 3 = 3 (π^+, π^-, π^0)

•Explicit symmetry breaking:

$$m_\pi^2 = B_0 \, m_q$$

gives mass to the NGB

Higgs is a pNGB



Physics Beyond the Standard Model 3.2

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Arequipa, Peru, March 6-19 2013

Beyond the Standard Model

Lecture I •Why do we need to go Beyond the SM ? •The Hierarchy Problem: what do we need to solve it ?

Lecture 2 • Supersymmetry and the Hierarchy Problem

Lecture 3 •New Dynamics at the TeV scale: the Higgs as a (pseudo) Nambu-Goldstone Boson

Beyond the Standard Model III.2

- Solve the Hierarchy problem with dynamics: QCD and the σ (Technicolor, ...)
- Dynamical (composite) <u>light</u> Higgs: is a (pseudo) Goldstone boson The example of the pion in QCD
 - •Composite Higgs Models:

Little Higgs Twin Higgs Gauge-Higgs unification in AdS_5

Higgs is a pseudo Nambu-Goldstone Boson Back to the analogy of QCD at low energies

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$$m_{\pi}^2 = B_0 \, m_q$$

gives mass to the NGB

Higgs is a pNGB



Higgs as a pNGB: Little Higgs Mechanism

•If the Higgs is (part of) a NGB then $m_h = 0$ and it can only have derivative interactions:

invariance under $h \rightarrow h + c$ shift symmetry

•In the SM $SU(2)_L \times U(1)_Y \to U(1)_{\rm EM}$

NGBs = # of broken generators = 3

But they're all eaten by W_L^{\pm}, Z_L

•To have NGBs left over need

Global symmetry > Gauge symmetry

Little Higgs Mechanism

- •The spontanous breaking of a global symmetry gives massless NGBS One of them: doublet of $SU(2)_L$ h
- But we need to give h a mass

 \Rightarrow Need explicit breaking of global symmetry

$$\implies m_h \neq 0$$

Try with global symmetry $SU(3) \rightarrow SU(2)$

- •# of broken generators = 8 3 = 5
- OK. We need 4 d.o.f. for h
- Explicit breaking to get $m_h \neq 0$ If we gauge part of SU(3) (e.g. SU(2)) we break explicitly SU(3) global \Longrightarrow quadratic divergences

If we gauge all of the SU(3)Global symmetry is respected. But now all of the NGBs are eaten to give masses to gauge bosons.

Solution: Enlarge the global symmetry to $SU(3) \times SU(3)$

 $SU(3) \times SU(3) \rightarrow SU(2) \times SU(2)$

of broken generators = 16 -6 = 10

But 5 eaten in the gauge breaking $SU(3) \rightarrow SU(2)$ So in the end: 5 NGBs left

 $\mathcal{L} = |D_{\mu}\Phi_{1}|^{2} + |D_{\mu}\Phi_{2}|^{2}$

$$\Phi_1 = e^{i\pi/f} \begin{pmatrix} 0\\ 0\\ f \end{pmatrix} , \Phi_2 = e^{-i\pi/f} \begin{pmatrix} 0\\ 0\\ f \end{pmatrix}$$

with

and

$$\pi = \pi^a t^a = \eta / \sqrt{2} + \begin{pmatrix} 0 & 0 & h_1 \\ 0 & 0 & h_2 \\ h_1^* & h_2^* & 0 \end{pmatrix}$$

$$\begin{pmatrix} h1\\ h2 \end{pmatrix} = h$$
 $SU(2)_L$ doublet

Gauge interactions do respect global symmetry

 $\mathcal{L} = |D_{\mu}\Phi_{1}|^{2} + |D_{\mu}\Phi_{2}|^{2}$

They do lead to quadratic divergences, from terms like

 $A_{\mu}A^{\mu}\left(\Phi_{1}^{\dagger}\Phi_{1}+\Phi_{1}^{\dagger}\Phi_{1}\right)$

$$\frac{g^2}{16\pi^2}\Lambda^2\left(\Phi_1^{\dagger}\Phi_1+\Phi_2^{\dagger}\Phi_2\right) = \frac{g^2}{16\pi^2}\Lambda^2\left(f^2+f^2\right)$$

But they do not induce and $h^{\dagger}h$ term Do not contribute to m_h^2

 \rightarrow

But at one loop we generate operators like $\Phi_1^{\dagger}\Phi_2$

$$\Phi_1 = e^{i\pi/f} \begin{pmatrix} 0\\ 0\\ f \end{pmatrix} , \Phi_2 = e^{-i\pi/f} \begin{pmatrix} 0\\ 0\\ f \end{pmatrix}$$

that will depend on h



Two boson propagators \implies only logarithmic divergence



But how are the quadratic divergences cancelled ?



Heavy gauge bosons cancel W quadratic divergence

Simplest Little Higgs: Fermions

We choose am anomaly-free model. Quarks:

$$\Psi_{Q_1} = \begin{pmatrix} d \\ u \\ D \end{pmatrix}_L \sim (\mathbf{3}^*, 0), \quad \Psi_{Q_2} = \begin{pmatrix} s \\ c \\ S \end{pmatrix}_L \sim (\mathbf{3}^*, 0),$$

$$\Psi_{Q_3} = \left(egin{array}{c} t \\ b \\ T \end{array}
ight)_L \sim (\mathbf{3}, 1/3),$$

and right-handed singlets

$$u_R, c_R, t_R \sim (1, 2/3), \quad d_R, s_R, b_R \sim (1, -1/3),$$

 $T_R \sim (1, 2/3)$ $D_R, S_R \sim (1, -1/3)$

Top Cancellation



Little Higgs

Electroweak Symmetry is broken radiatively

 $\delta V = \delta m^2 \, h^{\dagger} h + \lambda \, (h^{\dagger} h)^2$

Tension: get a light Higgs with large enough f

EWPC want $f \gtrsim 2 \text{ TeV}$

Many models other than this:

Littlest Higgs: SU(5)/SO(5) with $SU(2) \times SU(2)$

<u>T Parity</u>: Better agreement with EWPC, Dark Matter All require new fermions and new gauge bosons at f

Other pNGB Higgs Models

Twin Higgs Gauge-Higgs unification in AdS_5 . In all cases Higgs is composite Higgs couplings to SM particles is suppressed by powers of $\frac{v}{f}$

Separation of Scales in AdS₅

•One compact extra dimension. Non-trivial metric induces a small energy scale from a high one. (Randall, Sundrum '99)



• Geometry of extra dimension generates exponential hierarchy

$$\Lambda_{\rm TeV} \sim M_{\rm Planck} e^{-k L}$$
Separation of Scales in AdS5

• Warped 5D metric in RS

$$ds^{2} = e^{-2k|y|} \eta^{\mu\nu} dx_{\mu} dx_{\nu} - dy^{2}$$

- Compactified on orbifold S_1/Z_2 with $L=\pi R$ and k the AdS_5 curvature
 - •Hierarchy problem: for $kR \simeq (11 12)$

$$k e^{-k\pi R} \simeq O(1)$$
 TeV

The Hierarchy Problem in AdS

•If Higgs is localized in IR brane at $y = \pi R$

$$S_H = \int d^4x \int_0^{\pi R} dy \sqrt{-g} \,\delta(y - \pi R) \left[g_{\mu\nu} \partial^{\mu} H^{\dagger} \partial^{\nu} H - \lambda \left(|H|^2 - v_0^2 \right)^2 \right]$$

•Warp factor e^{ky} appears in $g_{\mu\nu}$ and $\sqrt{-g}$

$$S_H = \int d^4x \left[e^{-2k\pi R} \eta_{\mu\nu} \partial^{\mu} H^{\dagger} \partial^{\nu} H - e^{-4k\pi R} \lambda \left(|H|^2 - v_0^2 \right)^2 \right]$$

• Canonically normalize Higgs

$$e^{-k\pi R}H \to H$$

Hierarchy Problem in AdS₅

$$S_H = \int d^4x \, \left[\eta_{\mu\nu} \partial^{\mu} H^{\dagger} \partial^{\nu} H - \lambda \left(|H|^2 - e^{-2k\pi R} v_0^2 \right)^2 \right]$$

• If $v_0 \simeq M_P$, choosing $k R \simeq O(10)$ gives

 $v \simeq$ weak scale

\implies Higgs must be at or near IR brane

Bulk Fields in AdS

Bulk AdS models require

- •Enlarge gauge symmetry to include custodial symmetry in bulk to avoid large T parameter
- •Minimal choice: $SU(2)_L imes SU(2)_R imes U(1)_X$
- •Expand bulk theory in Kaluza-Klein modes. Get effective 4D theory

Gauge Fields in AdS Bulk

•Gauge fields zero-modes are flat by gauge invariance

•KK modes have IR-localized wave functions

 $M_n \simeq (n - O(1)) \times \pi \, k \, e^{-k\pi R}$

with masses starting at the TeV scale

Fermion Fields in AdS Bulk

•Massive fermion in curved 5D space

$$S_f = \int d^4x \, dy \, \sqrt{g} \left\{ \frac{i}{2} \bar{\Psi} \hat{\gamma}^M \left[\mathcal{D}_M - \overleftarrow{\mathcal{D}}_M \right] \Psi - M_f \bar{\Psi} \Psi \right\}$$

•To be natural $M_f \simeq O(1)k$

$$M_f \equiv c_f k$$
 with $c_f \simeq O(1)$

•The parameter c_f determines the localization of the ZM fermion

Fermion Fields in AdS Bulk

•E.g.: Localization of left-handed ZM

$$F_{\rm ZM}^L(y) = \frac{1}{\sqrt{2\pi R}} f_0^L(0) e^{(\frac{1}{2} - c_L) ky}$$

 $c_L > \frac{1}{2} \Rightarrow$ ZM fermion localized near Planck brane $c_L < \frac{1}{2} \Rightarrow$ ZM fermion localized near IR brane

Fermion Fields in AdS Bulk

O(I) flavor breaking in bulk can give fermion mass hierarchy



Fermions localized near TeV brane have O(I) Yukawas

Those localized near the Planck brane have highly suppressed Yukawas

Dynamical Localization of the Higgs

• Gauge-Higgs unification: gauge filed in 5D has scalar A_5

•To extract Higgs from A_5 need to enlarge gauge symmetry E.g.: $SU(3) \rightarrow SU(2) \times U(1)$ broken by boundary conditions

$$A_{\mu}: \begin{pmatrix} (+,+) & (+,+) & (-,-) \\ (+,+) & (+,+) & (-,-) \\ \hline (-,-) & (-,-) & (+,+) \\ \hline (-,-) & (-,-) & (+,+) \\ \hline (+,+) & (+,+) & (-,-) \end{pmatrix} \longrightarrow \text{Higgs doublet from} \quad A_{5} = A_{5}^{a} t^{a}$$

$$A_{5}: \begin{pmatrix} (-,-) & (-,-) & (+,+) \\ \hline (-,-) & (-,-) & (+,+) \\ \hline (+,+) & (+,+) & (-,-) \end{pmatrix}$$

Dynamical Localization of the Higgs

•In the dual 4D theory equivalent to Higgs as a NGB

gauge symmetry in bulk $A_5
ightarrow A_5 + \partial_y \chi$

 \Rightarrow shift symmetry in 4D

Higgs is a (p)NGB

The Flavor Problem in AdS₅

•KK gauge bosons couple stronger to heavier fermions



•Tree-level flavor violation is hierarchical. only important with heavier generations.

The Flavor Problem in AdS₅

- •Flavor bounds OK from most observables in K,D and B physics
- But one flavor observable is tough: ϵ_K mixed chirality operators $\bar{d}_R s_L \bar{d}_L s_R$

have large enhancement

$$\left(\frac{m_K}{m_s}\right)^2 \,\eta_1^{-5} \simeq 100$$

$$\Rightarrow M_{KK} > O(10) \text{ TeV}$$

•Requires flavor symmetries in the bulk

Conclusions

- Composite light Higgs require symmetry to protect m_h from being at the TeV scale (f)
- •Models of pNGB Higgs generally work require new global and gauge symmetries at fImply the existence of many new states above f
- •They also imply a new strong interaction above $4\pi f$
- •They replace a renormalizable theory with a non-renormalizable one ... but we've seen this before.